# Adaptive sequential Monte Carlo for posterior inference and model selection among complex geological priors

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### SUMMARY

Bayesian model selection enables comparison and ranking of conceptual subsurface models 2 described by spatial prior models, according to the support provided by available geophysз ical data. Deep generative neural networks can efficiently encode such complex spatial priors, thereby, allowing for a strong model dimensionality reduction that comes at the price of 5 enhanced non-linearity. In this setting, we explore a recent adaptive sequential Monte Carlo 6 (ASMC) approach that builds on Annealed Importance Sampling (AIS); a method that provides both the posterior probability density function (PDF) and the evidence (a central quan-8 tity for Bayesian model selection) through a particle approximation. Both techniques are well 9 suited to parallel computation and rely on importance sampling over a sequence of interme-10 diate distributions, linking the prior and the posterior PDF. Each subsequent distribution is 11 approximated by updating the particle weights and states, compared with the previous approx-12 imation, using a small pre-defined number of Markov chain Monte Carlo (MCMC) proposal 13 steps. Compared with AIS, the ASMC method adaptively tunes the tempering between neigh-14 boring distributions and performs resampling of particles when the variance of the particle 15 weights becomes too large. We evaluate ASMC using two different conceptual models and as-16 sociated synthetic cross-hole ground penetrating radar (GPR) tomography data. For the most 17 challenging test case, we find that the ASMC method is faster and more reliable in locating the 18

posterior PDF than state-of-the-art adaptive MCMC. The evidence estimates are found to be
 robust with respect to the choice of ASMC algorithmic variables and much less sensitive to the
 model proposal type than MCMC. The variance of the evidence estimates are best estimated
 by replication of ASMC runs, while approximations based on single runs provide comparable
 estimates when using a sufficient number of proposal steps in approximating each intermediate
 distribution.

Key words: Inverse theory, Statistical methods, Neural networks, Tomography, Ground pen etrating radar, Hydrogeophysics.

### 27 1 INTRODUCTION

Markov chain Monte Carlo (MCMC) methods are, for strongly non-linear inverse problems and a 28 limited computational budget, not always able to locate the posterior probability density function 29 (PDF) of interest or to explore it sufficiently. Parallel tempering (Earl & Deem, 2005) is a well-30 known approach to circumvent such issues and it was popularized in geophysics by Sambridge 31 (2014). Parallel tempering runs multiple interacting chains targeting a sequence of power posteri-32 ors including faster moving chains at higher temperatures (i.e., corresponding to less weight being 33 given to the likelihood function). Such chains help to locate significant modes of the posterior 34 distribution that can, through a swapping mechanism, be explored by the chain targeting the pos-35 terior distribution of interest for which the temperature is 1. The resulting increase in the ability to 36 bypass local minima and explore multimodal distributions is offset by the need for many parallel 37 chains and a carefully-tailored temperature sequence to ensure efficient mixing among chains. 38

<sup>39</sup> Neal (2001) introduced the annealed importance sampling (AIS) method, which is also well <sup>40</sup> suited to derive information about the posterior PDF of interest when confronted with highly non-<sup>41</sup> linear or multi-modal inverse problems. AIS is a particle method in which many particles (the <sup>42</sup> evolution of each particle is represented by an individual chain) are evolving in parallel. Particle <sup>43</sup> methods rely on the states and weights of a collection of evolving particles to approximate distri-<sup>44</sup> butions of interest. This is in contrast to MCMC methods in which all states have the same weight <sup>45</sup> and the distribution of interest is approximated by proposal and acceptance mechanisms ensuring

that sampling is proportional to the posterior probability density. In developing AIS, Neal (2001) 46 demonstrates how intermediate results obtained by simulated annealing (Kirkpatrick et al., 1983), 47 typically used for global optimisation, can be re-interpreted as a sequence of importance sampling 48 steps from approximations of intermediate posterior PDFs at gradually decreasing temperatures 49 (i.e., annealing), thereby, creating a succession of approximations of intermediate distributions 50 between the prior to the posterior distribution of interest. This method has several attractive prop-51 erties: (1) it inherits from simulated annealing the ability to bypass problems with local minima by 52 initially allowing large steps and efficient exploration before focusing on a more detailed search 53 in areas of high posterior probability; (2) it is well suited for parallelization; (3) the final states 54 and their associated importance weights approximate the posterior distribution; and (4) it offers 55 directly an approximation of the evidence, the central quantity in Bayesian model selection. 56

Even if AIS is still widely used, it suffers from two main deficiencies: (1) it is very challeng-57 ing to pre-define an appropriate annealing sequence (i.e., the sequence of inverse temperatures to 58 which the likelihood function is raised) and (2) the populations of importance weights have in-59 creasingly higher variances as the AIS run progresses, thereby, increasing the risk of obtaining 60 poor estimates of the posterior PDF and the evidence. Sequential Monte Carlo (SMC) (Doucet 61 & Johansen, 2011) represents a family of particle methods that are widely used in science and 62 engineering, particularly for data assimilation tasks, but their use in geophysics has been limited 63 to date (see review by Linde et al. (2017)). At the most basic level, SMC relies on importance 64 sampling combined with resampling steps which ensures that the particle approximation of the 65 high-dimensional posterior PDF is of sufficient quality. The resampling step tends to reinitialize 66 particles of low probability by states of higher probability, thereby avoiding that computational 67 time is wasted in regions of low posterior density. Zhou et al. (2016) proposed an adaptive SMC 68 algorithm (referred to hereafter as ASMC) that addresses the limitations of AIS stated above by 69 adaptively tuning the progression between intermediate distributions and by resampling when the 70 variance of the particle weights becomes too large. 71

The prior PDF has a strong impact on Bayesian geophysical inversion results (Hansen et al., 2012) and should reflect the existing geological knowledge at a site (see review by Linde et al.

(2015)). One effective way of encoding prior knowledge in a low-dimensional latent vector of 74 uncorrelated parameters is offered by deep generative neural networks (Goodfellow et al., 2014). 75 Laloy et al. (2017) and Laloy et al. (2018) demonstrated using variational autoencoders (Kingma 76 & Welling, 2013) and generative adversarial networks (GAN) (Goodfellow et al., 2014), respec-77 tively, that the generated realizations of such networks are of high quality and that inversion can 78 be successfully performed on this latent space. The challenge when working with deep generative 79 neural networks is the highly non-linear transform linking the latent variables to the image repre-80 sentation (i.e., the typically gridded model of physical properties). This non-linearity often leads 81 to poor and unreliable convergence when applying gradient-based optimization methods (Laloy 82 et al., 2019) and inversion on such latent spaces may challenge state-of-the-art MCMC algorithms 83 (Laloy et al., 2018). 84

Here, we explore the performance of the ASMC method (Zhou et al., 2016) when used together 85 with deep generative networks to approximate evidences and posterior distributions using geo-86 physical data. As examples, we consider crosshole geophysical ground-penetrating radar (GPR) 87 data and GAN-based priors, which implies highly non-linear and challenging inverse problems. In 88 ASMC, the approximations of intermediate posterior distributions is achieved by successively, at 89 each temperature, performing a small number of Markov steps. As model proposals, we consider 90 both an elaborate proposal scheme influenced by evolutionary algorithms and a basic uncorre-91 lated Gaussian proposal. Through these examples, we demonstrate that the ASMC method is: (1) 92 easy to implement in existing MCMC algorithms; (2) well-suited for parallelization; (3) robust 93 to parameter settings and model proposal schemes; (4) providing posterior approximations that 94 can be superior to those offered by state-of-the-art MCMC; and (5) deriving accurate evidence 95 estimations without strong distributional assumptions. 96

# 97 **2 METHOD**

In our method description below, we rely largely on the notation of Zhou et al. (2016) who intro duced the ASMC algorithm.

### **2.1** Bayesian inference and model comparison

Bayes' theorem expresses the posterior PDF of a conceptual model  $M_k$  with parameters  $\theta$ , given a set of observations y as:

$$\pi(\boldsymbol{\theta}|\mathbf{y}, M_k) = \frac{\pi(\boldsymbol{\theta}|M_k)p(\mathbf{y}|\boldsymbol{\theta}, M_k)}{\pi(\mathbf{y}|M_k)}.$$
(1)

<sup>103</sup> All the knowledge about the model parameters that is available before considering the data is <sup>104</sup> encapsulated in the prior PDF  $\pi(\theta|M_k)$ . The likelihood function  $p(\mathbf{y}|\theta, M_k)$  quantifies how likely <sup>105</sup> it is that a specific model realization gave rise to the observations when considering a prescribed <sup>106</sup> error model. The normalizing constant  $\pi(\mathbf{y}|M_k)$  is referred to as the evidence or the marginal <sup>107</sup> likelihood, and it is a multidimensional integral over the parameter space:

$$\pi(\mathbf{y}|M_k) = \int \pi(\boldsymbol{\theta}|M_k) p(\mathbf{y}|\boldsymbol{\theta}, M_k) d\boldsymbol{\theta}.$$
 (2)

The evidence quantifies the support provided by the data to the conceptual model under consideration, as formalized by the prior PDF, and can be used to rank different conceptual models. Schöniger et al. (2014) describe and compare different methods to estimate the evidence and found that numerical approaches generate more reliable estimates than mathematical approximations of equation 2 that yield analytical expressions. Recent studies comparing state-of-the-art approaches to evidence estimation in geophysical and hydrogeological contexts include Brunetti et al. (2017) and Brunetti et al. (2019).

## 115 2.2 Adaptive sequential Monte Carlo (ASMC)

### <sup>116</sup> 2.2.1 Importance sampling

Brute Force Monte Carlo (BFMC), also known as the arithmetic mean approach, evaluates many realizations drawn from the prior and the corresponding evidence estimate is their average likelihood. Unfortunately, BFMC suffers from the curse of dimensionality (Curtis & Lomax, 2001) in that most draws from the prior, when considering a handful or more unknown model parameters and high-quality data, have negligible likelihoods. Consequently, high likelihood regions contributing strongly to the mean are poorly sampled, leading to high-variance evidence estimates and
frequent underestimation of evidence values as demonstrated by Brunetti et al. (2017). Throughout
this manuscript, a high-variance estimate refers to that obtained by estimators of a mean quantity
(e.g., the mean of the sampled likelihoods) for which repeated estimations lead to widely different
estimates.

<sup>127</sup> Compared to BFMC, importance sampling offers lower-variance estimates, whereby Monte <sup>128</sup> Carlo samples are drawn proportionally to a so-called importance distribution  $q(\theta, M_k)$  (Hammer-<sup>129</sup> sley & Handscomb, 1964). In order to sample regions with a high contribution to the mean, this <sup>130</sup> distribution is chosen to be as close as possible to the target distribution; in this case the poste-<sup>131</sup> rior PDF. To account for the biased sampling procedure, every sample  $\theta^i$  drawn from  $q(\theta, M_k)$  is <sup>132</sup> associated with an importance weight defined as

$$w^{i} = \frac{\pi(\boldsymbol{\theta}^{i}|M_{k})p(\mathbf{y}|\boldsymbol{\theta}^{i},M_{k})}{q(\boldsymbol{\theta}^{i},M_{k})},$$
(3)

that determines the corresponding weight in the mean estimation. Assuming that  $q(\theta, M_k) \neq 0$ whenever  $\pi(\theta|M_k)p(\mathbf{y}|\theta, M_k) \neq 0$ , and if the number of draws  $N \to \infty$ , then the following approximation holds (Neal, 2001):

$$\frac{\sum_{i=1}^{N} w^{i}}{N} \approx \frac{\int \pi(\boldsymbol{\theta}|M_{k}) p(\mathbf{y}|\boldsymbol{\theta}, M_{k}) d\boldsymbol{\theta}}{\int q(\boldsymbol{\theta}, M_{k}) d\boldsymbol{\theta}}.$$
(4)

In the particular case of using the prior as the importance distribution (equivalent to BFMC) and noting that its evidence is equal to one (the integral of the prior PDF is 1), the evidence of  $M_k$ is approximated by the mean of the N weights:

$$\pi(\mathbf{y}|M_k) = \frac{\int \pi(\boldsymbol{\theta}|M_k) p(\mathbf{y}|\boldsymbol{\theta}, M_k) d\boldsymbol{\theta}}{\int \pi(\boldsymbol{\theta}, M_k) d\boldsymbol{\theta}} \approx \frac{\sum_{i=1}^N w^i}{N} = \frac{\sum_{i=1}^N \frac{\pi(\boldsymbol{\theta}^i|M_k) p(\mathbf{y}|\boldsymbol{\theta}^i, M_k)}{\pi(\boldsymbol{\theta}^i|M_k)}}{N} = \frac{\sum_{i=1}^N p(\mathbf{y}|\boldsymbol{\theta}^i, M_k)}{N},$$
(5)

which reduces to the average of the sampled likelihood as discussed above. The importance distri-

<sup>140</sup> bution strongly influences the accuracy of importance sampling and unreliable high-variance esti-<sup>141</sup> mates are obtained when the importance distribution is far from the target distribution. Therefore, <sup>142</sup> if the prior PDF is markedly different from the posterior PDF, then the quality of the evidence esti-<sup>143</sup> mate in equation 5 is low. Below, we explain how to obtain low-variance estimates of evidences by <sup>144</sup> relying on a succession of importance sampling steps with importance distributions that are close <sup>145</sup> to intermediate target distributions known as power posteriors.

### 146 2.2.2 Annealed importance sampling (AIS)

Simulated annealing (Kirkpatrick et al., 1983) is a well-known global optimizer that bypasses local 147 minima by gradually reducing the parameter space exploration using a sequence of intermediate 148 target distributions (i.e., power posteriors characterized by an annealing scheme of successively 149 decreasing temperatures). In developing AIS, Neal (2001) took advantage of this sequence of 150 transitional target distributions starting at the prior PDF (infinite temperature) and ending at the 151 posterior PDF (temperature of 1). The algorithm runs in parallel with each chain being interpreted 152 as a particle with an evolving weight and state. From the resulting sequence of intermediate im-153 portance weights and states, it is possible to estimate both the posterior PDF and the evidence. 154 AIS shares all the exploratory advantages of simulated annealing and allows for, potentially, high-155 quality posterior PDF and evidence estimations by creating a smooth path between the prior and 156 the posterior PDF. A schematic visualization of AIS is given in Figure 1a. 157

In the following, we consider a given conceptual model  $M_k$  and suppress the corresponding subindex k for simplicity. The unnormalized power posterior PDFs  $\{\gamma_t(\boldsymbol{\theta}_t | \mathbf{y})\}_{t=0}^T$  are:

$$\gamma_t(\boldsymbol{\theta}_t | \mathbf{y}) \equiv \pi(\boldsymbol{\theta}_t) p(\mathbf{y} | \boldsymbol{\theta}_t)^{\alpha_t}, \tag{6}$$

where  $\pi(\boldsymbol{\theta}_t)$  is the prior probability density function and  $p(\mathbf{y}|\boldsymbol{\theta}_t)$  the likelihood. The annealing schedule  $\alpha_t \in [0, 1]$  of inverse temperatures defines these power posteriors, where  $\alpha_{t=0} = 0$  gives the prior and  $\alpha_{t=T} = 1$  the posterior PDF. At small  $\alpha_t$ , the contribution of the likelihood is small and the corresponding power posterior is close to the prior PDF. As  $\alpha_t$  grows, the influence of



**Figure 1.** (a) Schematic representation of annealed importance sampling (AIS) using N = 4 particles evolving in parallel. Except for the initialization step, each color represents K = 4 Markov steps in which the particle system moves from approximating a previous unnormalized power-posterior to a new one. After each K = 4 Markov steps, the sampled states are used in an importance sampling step to determine the incremental weights  $w_t$  associated with the change in the intermediate posterior PDF. (b) In adaptive sequential Monte Carlo (ASMC), one main difference compared with AIS is that the  $\alpha$ -sequence determining the intermediate posterior distributions is no longer fixed but determined adaptively. Furthermore, resampling occurs when the variance of the weights are too large. Such a resampling step is here visualized with dashed red lines.

the likelihood on the power posterior increases. We denote  $Z_t$  as the normalizing constant of the corresponding power posterior, implying that the normalized power PDF is:

$$\pi_t(\boldsymbol{\theta}_t | \mathbf{y}) = \frac{\gamma_t(\boldsymbol{a}_t | \mathbf{y})}{Z_t}.$$
(7)

By using  $\gamma_{t-1}(\boldsymbol{\theta}_{t-1}|\mathbf{y})$  as an importance distribution for  $\gamma_t(\boldsymbol{\theta}_t|\mathbf{y})$ , we define the unnormalized incremental weights  $w_t$  for particle *i* at state  $\boldsymbol{\theta}_{t-1}^i$  as:

$$w_t^i = \frac{\gamma_t(\boldsymbol{\theta}_{t-1}^i|\mathbf{y})}{\gamma_{t-1}(\boldsymbol{\theta}_{t-1}^i|\mathbf{y})}.$$
(8)

Except for the initialization step, the corresponding importance distributions  $\gamma_{t-1}(\boldsymbol{\theta}_{t-1}|\mathbf{y})$  are approximated by updating N particles using K Markov steps targeting  $\gamma_{t-1}(\boldsymbol{\theta}_{t-1}|\mathbf{y})$  starting at a previous estimation of  $\gamma_{t-2}(\boldsymbol{\theta}_{t-2}|\mathbf{y})$ . Without these Markov steps, the AIS algorithm would reduce to BFMC. This process is schematized in Figure 1a for N = 4 and K = 4.

<sup>166</sup> It is customary to work with normalized weights defined as:

$$W_t^i = \frac{W_{t-1}^i w_t^i}{\sum_{j=1}^N W_{t-1}^j w_t^j},$$
(9)

where  $W_{t-1}$  are the previously defined normalized weights, that is,  $\sum_{i=1}^{N} W_{t-1}^{i} = 1$ . The final normalized weights  $W_{T}^{i}$  determine the relative probabilities of each of the final N states, thereby, approximating the posterior distribution through a particle approximation.

### 170 2.2.3 Evidence estimation

One major advantage of AIS and ASMC in the context of Bayesian model selection is that the evidence is readily obtained. The ratio of the normalizing constants of two consecutive intermediate distributions  $\gamma_t(\boldsymbol{\theta}_t | \mathbf{y})$  and  $\gamma_{t-1}(\boldsymbol{\theta}_{t-1} | \mathbf{y})$  is:

$$\frac{Z_t}{Z_{t-1}} = \frac{\int \gamma_t(\boldsymbol{\theta}_t | \mathbf{y}) d\boldsymbol{\theta}_t}{\int \gamma_{t-1}(\boldsymbol{\theta}_{t-1} | \mathbf{y}) d\boldsymbol{\theta}_{t-1}}, t$$
(10)

and it can be approximated as (Del Moral et al., 2006):

$$\frac{Z_t}{Z_{t-1}} \approx \sum_{i=1}^N W_{t-1}^i w_t^i.$$
 (11)

The posterior PDF of interest is the last distribution of the sequence ( $\alpha_{t=T} = 1$ ), therefore,

its normalizing constant is the evidence,  $Z_T = \pi(\mathbf{y})$ . Since the normalizing constant of the prior PDF,  $Z_0$ , is equal to one, the evidence can be estimated as the product of the intermediate ratios:

$$\pi(\mathbf{y}) = Z_T = \frac{Z_T}{Z_0} = \prod_{t=1}^T \frac{Z_t}{Z_{t-1}} \approx \prod_{t=1}^T \sum_{i=1}^N W_{t-1}^i w_t^i.$$
 (12)

### 178 2.2.4 Adaptive sequence of intermediate distributions

<sup>179</sup> Zhou et al. (2016) introduce several adaptations to AIS leading to the more robust ASMC algo-<sup>180</sup> rithm that requires much less tuning. The choice of the annealing schedule in equation (6) has <sup>181</sup> a strong impact on performance and it is generally difficult to assign a proper  $\alpha$ -sequence in ad-<sup>182</sup> vance. Zhou et al. (2016) solve this by introducing an adaptive procedure relying on the conditional <sup>183</sup> effective sample size (CESS):

$$CESS = N \frac{(\sum_{i=1}^{N} W_{t-1}^{i} w_{t}^{i})^{2}}{\sum_{j=1}^{N} W_{t-1}^{j} (w_{t}^{j})^{2}}.$$
(13)

The CESS measures the quality of the current intermediate distribution as an importance distribution to calculate expectations of the following one. To define the next distribution in the sequence (Figure 1a), a binary search is performed for the  $\alpha$ -increment for which the CESS is the closest to a pre-defined target value. The larger this target value is, the better the approximation, but the slower is the algorithm as the *L* number of intermediate distributions grows.

### 189 2.2.5 Resampling

The variance of the importance weights provides an indicator of the quality of the importance estimator. The importance weights invariably diverge over time leading to high variances, for example, because of poor convergence of some particles. To circumvent this, SMC methods rely on resampling (Del Moral et al., 2006; Doucet & Johansen, 2011). Resampling consists of reinitializing the states of each particle by replicating them according to a probability that is proportional to their current normalized weights. After resampling, the new states are assigned equal weights of 1/N. Figure 1b illustrates a resampling step. The purpose of this operation is to limit the variance

of the weights by excluding states with lower weights and replicating those with higher weights. 197 Since high-dimensional posterior distributions are estimated using N particles only, it is essential 198 that all samples contribute meaningfully to this approximation by avoiding regions of very low 199 probability. We rely herein on systematic resampling, which is easy to implement and performs 200 well with respect to alternative resampling schemes (Doucet & Johansen, 2011). The resampling 201 step impacts the variance of estimates (Douc & Cappe, 2005) and it is often beneficial to only per-202 form resampling occasionally. To decide when to apply resampling, we follow standard practice 203 by relying on a quantity that considers the history of the weight variance evolution, namely the 204 effective sample size (ESS) (Kong et al., 1994): 205

$$ESS_t = \frac{(\sum_{i=1}^{N} W_{t-1}^i w_t^i)^2}{\sum_{j=1}^{N} (W_{t-1}^j)^2 (w_t^j)^2}.$$
(14)

The ESS can be interpreted as reflecting the number of effective samples in the particle approximation and resampling is applied when the ESS is lower than a pre-defined threshold.

### 208 2.2.6 Evidence uncertainty estimation

The most reliable approach to assess uncertainty on evidence estimates is to perform multiple 209 ASMC runs and calculate the resulting variance of the estimates. This is the approach used by Zhou 210 et al. (2016) when introducing ASMC. Even if such Monte Carlo replication is easily parallelized, 211 it implies a significant computational overhead as the total computational effort grows linearly with 212 the number of replicates. In recent years, progress has been made in obtaining evidence variance 213 estimates from single SMC runs. The first consistent estimator was proposed by Chan & Lai (2013) 214 and a refined estimator was later introduced by Lee & Whiteley (2018). We consider a modified 215 form of this latter estimator in Doucet & Lee (2018) that we adopted to account for occasional 216

resampling. The resulting expression should be interpreted as a relative variance contribution of
the evidence estimate contribution since the last resampling time:

$$\frac{V_t^N}{(\eta_t^N)^2} = \frac{1}{(\eta_t^N)^2} \left(\frac{N}{N-1}\right)^n \frac{1}{N(N-1)} \sum_{i=1}^N \left[\sum_{j:E_t^j=i} (NW_{t-1}^j w_t^j - \eta_t^N)\right]^2,$$
(15)

where  $\eta_t^N = \sum_{i=1}^N NW_{t-1}^i w_t^i$  and n is the cumulative number of resampling steps that has been performed until t. The index  $E_t^j$  is the so-called Eve index of particle j at time t, which traces the origin of the particles. If no resampling is done, the Eve indices stay constant and are equal to 1 : N. After resampling, the states of the particles are reorganized and the Eve indices change, denoting the original particle that moved to that position. A graphical illustration of this process is given by Lee & Whiteley (2018). The number of remaining unique Eve indices along the run can be interpreted as a conservative estimate of the number of independent particles.

We compute the estimator in equation 15 before each resampling step and at the last step of the ASMC algorithm. We then sum the resulting contributions:

$$\sigma_r = \sqrt{\sum_{h=0}^R \frac{V_h^N}{(\eta_h^N)^2}},\tag{16}$$

where R is the total number of resampling times. This equation is valid under the assumption that the individual contributions in the sum are independent (Brown & Neal, 1991). Hence, we assume here that the particles decorrelate from each other between resampling steps.

### 231 2.2.7 Markov proposals and acceptance criteria

We implemented ASMC within the popular Differential Evolution Adaptive Metropolis ZS (DREAM<sub>(ZS)</sub>) algorithm (Laloy & Vrugt, 2012). In this MCMC algorithm, model proposal updates with respect to the present state are drawn proportionally to random differences of past states, thus, helping to better explore the target distribution by automatically determining the scale and direction of the model proposals. If we consider J as a  $m \times d$  dimensional matrix that contains m past states of the chains, where d is the number of parameters, the jump vector for the *i*-th chain is given by (Vrugt, 2016):

$$d\boldsymbol{\theta}_{A}^{i} = \zeta_{d^{*}} + (1_{d^{*}} + \lambda_{d^{*}})\psi(\delta, d^{*}) \sum_{j=1}^{\delta} (\mathbf{J}_{A}^{a_{j}} - \mathbf{J}_{A}^{b_{j}}).$$
(17)

If the current state is  $\theta^i$ , then the candidate point for particle *i* is  $\theta^i_{prop} = \theta^i + d\theta^i$ . The num-239 ber of pairs used to generate the jump is given by  $\delta$ , and a and b are vectors of integers drawn 240 without replacement from  $\{1, ..., m\}$ . The parameters  $\zeta$  and  $\lambda$  are sampled independently from 241 pre-defined uniform and normal distributions, respectively. This algorithm implements subspace 242 sampling, which implies that only a random subset A of  $d^*$ -dimensions from the original param-243 eter space is updated at each proposal step. The difference between past states is multiplied by a 244 fixed proposal scale referred to as jump rate  $\psi(\delta, d^*) = \frac{2.38}{\sqrt{2\delta d^*}} \epsilon$ , where  $\epsilon$  is an user-defined factor 245 that we introduce to further control the size of the jumps. In contrast to MCMC, ASMC allows 246 straightforward adaptation of the  $\epsilon$ -factor on-the-go without violating detailed balance condition. 247 This tuning of  $\epsilon$  is achieved by using the acceptance rate (AR) of the last K Markov steps to target 248 an acceptance rate above  $AR_{min}$ . To implement this,  $\epsilon$  is initialized to a comparatively large value 249 and a percentage decrease of its value f is made when the acceptance rate falls below  $AR_{min}$ . For 250 comparison purposes, we also consider standard model proposals given by uncorrelated Gaussian 251 draws centered on the previous state. For this case, the jump vector for the *i*-th chain is given by: 252

$$d\boldsymbol{\theta}_A^i \stackrel{i.i.d.}{\sim} \mathcal{N}_A(0, \epsilon^2). \tag{18}$$

Our considered model proposals are symmetric and the prior PDF is uniform. Consequently, with proper boundary handling, the proposed moves are accepted according to the likelihood ratio (Mosegaard & Tarantola, 1995). The probability to accept each candidate model during the *K* Markov steps used to approximate  $\gamma_t(\boldsymbol{\theta}_t|y)$  is:

$$P = \min\left\{1, \frac{p(\mathbf{y}|\boldsymbol{\theta}_{prop})^{\alpha_t}}{p(\mathbf{y}|\boldsymbol{\theta})^{\alpha_t}}\right\}.$$
(19)

Assignment of user-defined variables: Define number of particles (N), optimal CESS ( $CESS_{op}$ ), ESS threshold ( $ESS^*$ ), number of MCMC iterations at each intermediate distribution (K), minimal acceptance rate ( $AR_{min}$ ), initial proposal scale factor ( $\epsilon$ ) and its percentage decrease (f). Initialization: Set t = 0Set  $\alpha = 0$ Sample  $\theta_0$  from the prior  $\pi(\theta_t | M_k) N$  times Set the N-dimensional vector of normalized weights  $\mathbf{W}_0 = [\frac{1}{N}; \frac{1}{N}; ...; \frac{1}{N}]$ Set evidence  $\pi(\mathbf{y}|M_k) = 1$ Iteration : Set t = t + 1Search for incremental distribution Do binary search for the increment  $\Delta \alpha$  that gives the CESS (eq. 13) that is the closest to  $CESS_{op}$ . Update  $\alpha = min(1, \alpha + \Delta \alpha)$  and define the intermediate distribution  $\gamma_t(\boldsymbol{\theta}_t | \mathbf{y}) = \pi(\boldsymbol{\theta}_t | M_k) p(\mathbf{y} | \boldsymbol{\theta}_t)^{\alpha}$ . Compute the weight increments  $w_t^i$  (eq. 8), update and save the normalized weights  $W_t^i$  (eq. 9) and the evidence  $\pi(\mathbf{y}|M_k) = \pi(\mathbf{y}|M_k) \sum_{i=1}^N W_{t-1}^i w_t^i$  (eq.12). Resampling Calculate ESS (eq. 14), if  $ESS < ESS^*$  do resampling: re-organize  $\theta_t$  states and update  $\mathbf{W}_t = [\frac{1}{N}; \frac{1}{N}; ...; \frac{1}{N}]$ Do K MCMC iterations for each of the N particles (chains): Propose moves  $\theta_{prop}$  (eq. 17 and 18) and accept or reject based on acceptance criterion (eq. 19) using  $\gamma_t(\boldsymbol{\theta}_t | \mathbf{y})$ . Save the  $N \theta$  and their likelihoods. Set last state as  $\theta_{t+1}$ Tune proposal scale

### Algorithm 1: ASCM algorithm adopted from Zhou et al. (2016); their algorithm 4.

#### 2.2.8 Full ASMC algorithm 257

Repeat until  $\alpha = 1$ 

The full algorithm is given in Algorithm 1, for which the total number of iterations per consid-258 ered particle (chain) is equal to L (number of intermediate distributions)  $\times K$  (MCMC steps per 259 distribution). 260 This algorithm has several important strengths: (i) it requires a rather small number of user-261

If acceptance rate  $AR < AR_{min}$  then decrease proposal scale factor:  $\epsilon = \epsilon * (1 - \frac{f}{100})$ 

defined parameters; (ii) the posterior PDF and the evidence are estimated; (iii) the variance of the 262

weights are used to assess accuracy, (iv) the adaptation of classical MCMC algorithms into ASMC 263

is straightforward, and (v) the acceptance rate is controlled throughout the inversion. 264

#### The Laplace-Metropolis method 2.3 265

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MCMC algorithms provide an approximation of the posterior distribution, however, they need 266 to be combined with an additional estimation procedure to provide evidence estimates. For later 267 comparison purposes with ASMC, we mention here the Laplace-Metropolis estimator (Lewis & 268 Raftery, 1997), a mathematical approximation of the evidence using a Taylor expansion around the maximum a posteriori (MAP) estimate. Assuming that the posterior PDF is well approximated
by a normal distribution, the resulting evidence estimate is:

$$\pi(\mathbf{y}|M_k) = (2\pi)^{\frac{d}{2}} |\mathbf{H}(\boldsymbol{\theta}^*)|^{\frac{1}{2}} \pi(\boldsymbol{\theta}^*|M_k) p(\mathbf{y}|\boldsymbol{\theta}^*, M_k),$$
(20)

where  $\theta^*$  is the MAP estimate, *d* is the number of parameters and  $|\mathbf{H}(\theta^*)|$  is the determinant of minus the inverse Hessian matrix evaluated at the MAP, which is approximated from the MCMCbased samples from the posterior.

### **275** 2.4 From implicit to prescribed geostatistical priors

Multiple-point statistics (MPS) (Mariethoz & Caers, 2014) is a sub-field of geostatistics aiming at 276 producing conditional geostatistical model realizations of high geological realism, thereby, cap-277 turing more meaningful connectivity patterns than those offered, for instance, by classical mul-278 tivariate Gaussian priors (Renard & Allard, 2013). MPS algorithms produce model realizations 279 that are in agreement with the spatial patterns found in a so-called training image (TI). A TI is a 280 gridded representation of the targeted spatial field obtained from geological information such as 281 outcrops or process-based simulation methods (Koltermann & Gorelick, 1996). Performing inver-282 sion (Mariethoz et al., 2010; Hansen et al., 2012; Linde et al., 2015) and model selection (Brunetti 283 et al., 2019) based on one or more TIs commonly requires inversion algorithms that work with 284 so-called implicit priors. That is, the MPS algorithm provides model realizations that are drawn 285 proportionally to the prior, but the prior density of a given realization is unknown. Two main issues 286 arise with this approach: (1) the generation of conditional prior realizations may be computation-287 ally expensive in MCMC settings when a large number of model proposals are needed, and (2) 288 the implicit prior model precludes the calculation of prior probability densities as needed in many 289 state-of-the-art inversion and model selection methods. 290

Deep learning (LeCun et al., 2015) applied to geoscientific problems has been growing rapidly in recent years (Bergen et al., 2019; Karpatne et al., 2018). In particular, deep generative neural networks offer an attractive approach to build an explicit prior PDF from training images (Laloy

et al., 2017, 2018; Mosser et al., 2017, 2020), that is, a prior for which the prior density of any 294 realization is easily calculated. This is achieved by learning a non-linear transform between a low-295 dimensional latent space with a prescribed prior (typically an uncorrelated standard normal or 296 bounded uniform prior) and the image space (on which the forward simulations are performed). 297 To do this, the neural network is trained repeatedly with pieces of a large TI or MPS realiza-298 tions. Such tailor-made model parametrizations achieve significant dimensionality reduction by 299 leveraging spatial patterns in the TI. Inversion is then performed on the latent space and the re-300 sulting posterior is mapped, using the trained transform, into a posterior on the original image 301 space (a so-called push-forward operation). We rely on a spatial generative adversarial neural net-302 work (SGAN) (Jetchev et al., 2016), where each dimension of the latent space influences a given 303 region of the generated image space. The network's weights are learned by adversarial training 304 (Goodfellow et al., 2014). The latter consists of a competition between a so-called discriminator 305 and a generator: the discriminator aims to distinguish fake (i.e., realizations by the generator) and 306 real (i.e., training samples) images, while the generator tries to fool it by generating realizations 307 similar to the training samples. This is mathematically translated in a minimization-maximization 308 problem (see the book by Goodfellow et al. (2016), for details). The main computational effort 309 is related to training and once trained, the computational cost to draw model proposals in the la-310 tent space and to map them into the image space (for further forward computations) is very low. 311 The motivation of evaluating ASMC using a deep-learning based parameterization is two-fold: 312 (1) the SGAN parameterization implies strong non-linearity which makes it difficult for MCMC 313 algorithms to converge when performing inversion on the SGAN latent space (Laloy et al., 2018), 314 thus providing challenging test examples for which the added value of ASMC for posterior infer-315 ence can be demonstrated and (2) to build on recent work (Brunetti et al., 2019) on MPS-based 316 Bayesian model selection to highlight the value of prescribed priors when performing model se-317 lection among MPS-based prior models. 318

16

### 319 3 RESULTS

### 320 **3.1 Test examples**

Two conceptual 2-D models represented by TIs were used to assess ASMC for inversion and 321 model selection purposes. These TIs are used to train SGANs that generate realizations honoring 322 the multiple-point statistics of the TIs (Laloy et al., 2018). The first conceptual model (Figure 323 2a) is represented by a binary channelized training image (CM1) (Zahner et al., 2016) and the 324 second one (Figure 2b) is represented by a tri-categorical training image characterizing braided-325 river aquifer deposits (CM2) (Pirot et al., 2015). The SGAN generators are assigned uniform priors 326 on the latent space: the CM1-realizations and the more complex CM2-realizations have 15 and 45 327 latent variables, respectively. All realizations correspond to an image dimension of  $129 \times 65$  cells 328 that is cropped to  $125 \times 60$ , with a discretization of 0.1 m  $\times$  0.1 m (Figure 3). 329

Our synthetic data correspond to simulated crosshole ground-penetrating radar (GPR) first-330 arrival travel times with a geometry consisting of two boreholes that are 5.8 m apart. A total of 331 24 sources and 24 receivers are placed equidistantly every 0.5 meters in depth. First-arrival times 332 were calculated using the *time2d* algorithm by Podvin & Lecomte (1991). Following common 333 practice, the data were filtered according to a maximum angle between sources and receivers of 334 45 degrees (Peterson, 2001), resulting in 444 travel times. In order to assign velocities to each 335 facies, the corresponding dielectric constants were approximated using the complex refractive 336 index method (CRIM) (Roth et al., 1990). Representative porosities for CM2 were taken from 337 Pirot et al. (2019) and adjusted to CM1 to have the same mean and variance. The two reference 338 models used to produce the synthetic data are shown in Figure 3. They were obtained as a randomly 339 chosen realization from the respective SGAN generators. Uncorrelated Gaussian random noise 340 with standard deviation  $\sigma = 1$  ns was added to the resulting travel times simulated from these 341 models. 342

### **343 3.2 ASMC performance**

We first present the parameter settings and the performance of the ASMC algorithm (section 2.2.8) using DREAM<sub>(ZS)</sub> proposals (ASMC-DREAM) with N = 40 particles. To tune the proposal scale,



**Figure 2.** Training images: (a)  $2500 \times 2500$  binary channelized training image (CM1) (Zahner et al., 2016) and (b)  $400 \times 400$  tri-categorical training image representing a braided aquifer (CM2) (Pirot et al., 2015). The discretization of the cells is 0.1 m × 0.1 m.

we apply a 20% decrease (f = 0.2) with  $AR_{min} = 0.25$ . The starting large proposal scale  $\epsilon$  is gradually decreased as the annealing progresses (i.e., the inverse temperature  $\alpha$  increases towards 1). We implemented adaptive selection of the  $\alpha$ -sequence, using a binary search defined on a range



**Figure 3.** Reference models with associated velocities. (a) CM1: channel velocity v=0.085 m/ns and matrix velocity v=0.071 m/ns. (b) From Pirot et al. (2015) CM2: gray gravel (gray) v=0.083 m/ns, open framework (black) v=0.065 m/ns and bimodal (white) v=0.086 m/ns. Red stars and blue triangles represent GPR sources and receivers, respectively.

of  $\alpha$ -increments from  $10^{-5}$  to  $10^{-2}$ , to find the increments with the *CESS* that is the closest to the target  $CESS_{op}$ . The  $CESS_{op}/N$  ratio is in practice chosen close to 1. The closer it is to 1, the higher the number of intermediate distributions and the larger is the quality of estimates. Resampling is applied whenever ESS/N falls below 0.5. Table 1 contains the user-defined parameters and the resulting sequence lengths. The total number of forward simulations of each ASMC run is  $N \times K \times L$ .

Figures 4(a-b) show the evolution of the likelihood raised to the power of the corresponding 355  $\alpha$  in the natural log-scale for CM1 and CM2, respectively. This type of plotting is consistent 356 with the target distribution  $\gamma_t(\boldsymbol{\theta}_t | \mathbf{y})$  at each step (equation 6). The black dashed line indicates the 357 target log-likelihood calculated with the random noise realization used to noise-contaminate the 358 forward response of the reference model, raised to the power of the corresponding  $\alpha$ . Figures 4(c-d) 359 present correspondingly the acceptance rate evolution. As  $\alpha$  grows, the acceptance rate for a given 360 jump rate decreases as the targeted posterior distribution gives larger weights to the likelihood. 361 When the acceptance rate falls below  $AR_{min} = 0.25$ , the proposal scale is reduced causing a 362 small increase, after which the acceptance rate starts decreasing again until another reduction of 363 the proposal scale is required, thereby, keeping the acceptance rate in a range between 25% and 364 40%. Figures 4e-f show the optimized sequence of  $\alpha$ -values defining the intermediate posterior 365 distributions, obtained through a binary search of the  $\alpha$ -increments. In Figures 4g-h, the logarithm 366 of the normalized weight of each particle is plotted against the  $\alpha$ -index. Finally, Figures 4i-j shows 367 the evolution of the natural logarithm of the evidence vs.  $\alpha$ . 368

To ensure convergence with the more complex test case CM2, we had to choose a higher 369 CESS<sub>op</sub> and K than for CM1, which resulted in an approximately 4.7 times longer run. Despite 370 these adaptations, more resampling steps were needed compared to CM1 (see Table 1), which 371 reinforces the impression that it is a more challenging scenario. The increasing complexity of 372 CM2 is also indicated by the fact that the intermediate target distributions are well-approximated 373 for CM1 (Figure 4a) for which the sampled likelihoods fall close to the dashed line, while this is 374 less the case for CM2 (Figure 4b). However, both test cases reached the target log-likelihood and 375 the resampling fulfills its role of limiting the variance of the weights. 376



**Figure 4.** Results of ASMC with DREAM<sub>(ZS)</sub> model proposals (ASMC-DREAM) for conceptual models CM1 (left column) and CM2 (right column): (a) and (b) natural logarithm of the likelihood to the power of  $\alpha$  vs. iterations per particle. Each color represents a different particle and the black dashed line indicates the logarithm of the likelihood to the power of  $\alpha$  calculated using the random noise realization used to noise-contaminate the forward-simulated true model; (c) and (d) acceptance rate vs. iterations per particle, the dashed line indicates a 25% threshold; (e) and (f)  $\alpha$ -sequence vs.  $\alpha$  index; (h) and (i) natural lognormalized weights vs.  $\alpha$  index where each color represents a different particle; (j) and (k) natural log-evidence evolution vs.  $\alpha$ .

Algorithm 1 is applicable to other model proposals than  $DREAM_{(ZS)}$ . This is demonstrated using standard (vanilla) MCMC model proposals based on uncorrelated random Gaussian perturbations (ASMC-Gauss). In this case, the algorithm starts with a high standard deviation of the centered Gaussian model proposal and it is subsequently decreased when the acceptance rate falls
below 25%. The user-defined parameters were chosen to be the same as for the ASMC-DREAM
tests detailed in Table 1, leading to a similar sequence length as for ASMC-DREAM. The corresponding results are shown in Figure 5. For CM1, ASMC-Gauss needed one more resampling
time (Fig. 5c) compared to ASMC-DREAM due to a faster increase in the variance of the weights.
Otherwise, the performance of ASMC-DREAM (Figure 4) and ASMC-Gauss (Figure 5) are very
similar.

### **387 3.3** MCMC performance

For comparative purposes, we also perform MCMC inversions (no ASMC) using 40 chains and 388 a similar number of forward simulations. Again, we consider two tests: one using DREAM<sub>(ZS)</sub> 389 (MCMC-DREAM) and one with random Gaussian perturbations (MCMC-Gauss). Extensive man-390 ual tuning of the inversion parameters was needed to achieve satisfactory results. Figure 6 shows 391 the results obtained for conceptual models CM1 and CM2. The log-likelihood evolution is shown 392 in Figures 6a-d and the acceptance rate in Figures 6e-h. In order to assess convergence, the po-393 tential scale reduction factor  $\hat{R}$  is calculated (Gelman & Rubin, 1992) and plotted in Figures 6i-l, 394 with convergence declared when  $\hat{R}$  is below 1.2 for all model parameters. 395

The only MCMC run reaching convergence is MCMC-DREAM for CM1 at around 10,000 396 iterations. For this conceptual model, the results obtained with MCMC-Gauss are unsatisfactory 397 with only a few of the chains approaching the target likelihood, while the others are trapped in lo-398 cal minima, thereby, demonstrating a vastly superior performance of MCMC-DREAM compared 399 with MCMC-Gauss. For CM2, none of the MCMC inversions converge within the allotted com-400 putational time, as  $\hat{R}$  does not fall below 1.2. This is also reflected in the likelihood evolution: the 401 majority of sampled likelihoods remains below the target likelihood along the run. To summarize, 402 we find for a similar computational budget that the ASMC algorithm reaches the target likelihood 403 for both conceptual models and model proposal types, while the MCMC runs only approximate 404 the target likelihood for CM1 using MCMC-DREAM. 405



**Figure 5.** Results of ASMC with standard MCMC (ASMC-Gauss) for conceptual models CM1 (left column) and CM2 (right column): (a) and (b) natural logarithm of the likelihood to the power of  $\alpha$  vs. iterations per particle. Each color represents a different particle and the black dashed line indicates the logarithm of the likelihood to the power of  $\alpha$  calculated using the random noise realization used to noise-contaminate the forward-simulated true model; (c) and (d) acceptance rate vs. iterations per particle, the dashed line indicates a 25% threshold; (e) and (f)  $\alpha$ -sequence vs.  $\alpha$  index; (h) and (i) natural log-normalized weights vs.  $\alpha$  index, each color represents a different particle; (j) and (k) natural log-evidence evolution vs.  $\alpha$ .

### **3.4 Posterior distributions**

- <sup>407</sup> We focus now on the posterior approximations obtained with ASMC-DREAM and MCMC-DREAM.
- <sup>408</sup> For MCMC-DREAM, the posterior is obtained by first removing the so-called burn-in period, that



**Figure 6.** MCMC inversion results from DREAM<sub>(ZS)</sub> (MCMC-DREAM) and standard MCMC with Gaussian model proposals (MCMC-Gauss) for conceptual models CM1 and CM2. (a)-(d) the natural logarithm of the likelihood vs. iterations, where each color represents a different particle and the black dashed line indicates the log-likelihood calculated using the random noise realization, (e)-(h) the acceptance rate evolution, and (i)-(l) the evolution of the potential scale reduction factor  $\hat{R}$  with each color representing a different parameter and the black dashed lines indicating the value below which convergence is declared ( $\hat{R} = 1.2$ ).

is, the number of iterations needed to reach the target likelihood, from which it starts to sample
from the posterior PDF. The remaining samples contribute equally to the posterior estimations.
This is not the case for ASMC, for which the posterior PDF is approximated by the last states and

- <sup>412</sup> weights of the particles (chains).
- For a smoother representation of the posterior PDF approximated by ASMC, we applied kernel

 Table 1. ASMC user-defined parameters and resulting sequence length for conceptual models CM1 and CM2.

	ASMC-DREAM CM1	ASMC-DREAM CM2	ASMC-Gauss CM1	ASMC-Gauss CM2
Particles (N)	40	40	40	40
$CESS_{op}/N$	0.999993	0.999996	0.999993	0.999996
$ESS^*/N$	0.5	0.5	0.5	0.5
$AR_{min}$	25%	25%	25%	25%
K iterations	20	60	20	60
L intermediate distributions	4798	7775	4871	7673
Iterations per particle	95960	466500	97420	460380
Resampling times	1	5	2	3
Total numerical demand [ $\times 10^5$ ]	38.384	186.600	38.968	184.152

density estimation (KDE) (Scott, 2015). Figure 7 compares the estimated posteriors for CM1. The
KDE bandwidth impacts on the level of smoothing, that we chose to kept fixed for the 15 parameter
posteriors. Nevertheless, the estimated posteriors are overall very similar, which suggests that
ASMC provides a good estimation of the posterior. No comparison is provided for CM2 as the
MCMC-DREAM algorithm did not converge, neither in terms of reaching the target likelihood
nor in terms of exploration of the posterior PDF.

We now consider the posterior means and variances in the image space by translating the poste-420 rior realizations in the latent space using the SGAN generator. For ASMC-DREAM, the mean and 421 standard deviation images correspond to the last states of the chains weighted by their weights. 422 For MCMC-DREAM, the mean and standard deviation images are obtained using the equally 423 weighted states in the second half of the chains. The means and standard deviations for CM1 are 424 very similar for ASMC-DREAM (Figure 8b-c) and MCMC-DREAM (Figure 8d-e) that both ap-425 proximate the true model very well (Figure 8a). For CM2, we see a much better defined mean 426 model and smaller standard deviations for ASMC-DREAM (Figure 8g-h). The poorer approxima-427 tions by MCMC-DREAM 8i-h) is a direct consequence of the fact that this run did not converge. 428 Table 2 shows the log-likelihood range for the different inversions. For MCMC-DREAM, the sec-429 ond halves of the chains are considered for the range, while only the last states of the particles are 430 considered for ASMC-DREAM. 431

### **432 3.5** Evidence estimation

Even if the theoretical basis of the ASMC method for evidence estimation is well-established (Zhou et al., 2016), we start this section by considering a simple example that allows for comparison with BFMC (see section 2.2.1). We consider CM1 in a high-noise setting using uncorrelated Gaussian random noise with standard deviation  $\sigma = 15$  ns. This is certainly an unrealistically high noise level, but it allows us to obtain reliable evidence estimates through BFMC using 2 million prior samples. The resulting log-evidence obtained by BFMC is -1798.92, while the corresponding ASMC-DREAM run using K = 5 and  $CESS_{op}/N = 0.9999$  (resulting in 1100 iterations per



**Figure 7.** Estimated marginal posterior distributions for CM1 using ASMC with  $DREAM_{(ZS)}$ -proposal (ASMC-DREAM) and regular  $DREAM_{(ZS)}$  (MCMC-DREAM) with a comparable number of forward computations. Results are shown for all latent model parameters that have bounded uniform priors between -1 and 1.

particle) led to a log-evidence estimate of -1798.86, which is practically identical to the BFMC
 estimate.

After having established that our ASMC implementation provides accurate evidence estima-442 tion by comparison with BFMC, we now return to the original low-noise  $\sigma = 1$  ns setting. For the 443 test examples considered in the previous sections, the evidence estimates obtained with ASMC-444 DREAM and ASMC-Gauss given in Table 2 (i.e., the last computed values shown in Figures 4i-j 445 and 5i-j) are very close to each other. For comparison purposes, we also calculate the Laplace-446 Metropolis evidence estimator (LM) using the MCMC-DREAM inversion results (equation 20). 447 This is done for CM1 only as MCMC-DREAM did not converge for CM2. The Laplace-Metropolis 448 estimate (Table 2) is only slightly lower than the ASMC-DREAM and ASMC-Gauss estimates. 449 The close agreement between ASMC-DREAM and ASMC-Gauss, and the close agreement con-450



**Figure 8.** Reference model for (a) CM1 and (f) CM2; mean of the weighted final states from ASMC-DREAM for (b) CM1 and (g) CM2; standard deviations of the corresponding weighted final states for (c) CM1 and (h) CM2; mean of the second half of the MCMC chains obtained with MCMC-DREAM for (d) CM1 and (i) CM2 (not converged); corresponding standard deviations for (e) CM1 and (j) CM2 (not converged).

sidering the simplifying assumptions of the Laplace-Metropolis method, suggest again that the
 results obtained with ASMC are accurate.

Until now, we have considered that the right conceptual (prior) model was used in the inver-453 sions. That is, the noise-contaminated data were generated with a realization of the assumed prior 454 PDF. We now consider how the evidence changes if we make the wrong assumption, that is, use 455 the noise-contaminated data generated from a prior draw of another conceptual model. In Figure 456 9 we display the evidence evolution for two such incorrect scenarios using ASMC-DREAM with 457 combinations of CM1 and CM2 in the data generation and inversion process. The resulting log-458 evidence estimates (Table 2) are many hundreds of times smaller than the estimations obtained by 459 making the right assumption, suggesting in these simple scenarios that the true conceptual model 460 can easily be inferred if it is in the set of considered conceptual models. 461

### **3.6** Evidence uncertainty quantification

<sup>463</sup> We first assess the uncertainty of the evidence estimations by performing Monte Carlo replication.

<sup>464</sup> For the low noise ASMC-DREAM tests shown in section 3.2, we performed ten separate runs of



**Figure 9.** ASMC-DREAM evidence evolution with respect to the  $\alpha$ -sequence evolution when making incorrect assumptions about the underlying conceptual model: (a) CM1-based prior in the inversion using data generated from a prior realization from CM2, and (b) CM2-based prior in the inversion using data generated from a prior realization from CM1.

ASMC-DREAM for CM1 and five for CM2. We varied K and kept all other parameters fixed.

<sup>466</sup> Figure 10 shows the corresponding evidence estimations for CM1 and their means in logarithmic

467 units. Table 3 shows the relative standard deviation for both conceptual models. For CM1, it de-

creases almost by a factor of 10 when moving from K = 1 to K = 20. For this case, even K = 1

<sup>469</sup> leads to rather high-quality estimates with a relative standard deviation of 1.72. The decrease is

 $_{470}$  less abrupt for CM2 when increasing K = 5 to K = 60.

From a computational standpoint, it is beneficial if high-quality uncertainty estimates of the evidences would be obtained from one ASMC run only. Hence, we assess how the predictions of equations 15 and 16 compare with the estimates based on Monte Carlo replications. For smaller K, resampling compensates for the faster increasing variance of the weights, but this is at the expense

**Table 2.** Natural log-likelihood range, natural log-evidence estimation and number of resampling steps for the different inversion cases. The log-likelihoods of the reference models are -642.34 (CM1) and -616.00 (CM2).

	Log-likelihood range	Log-evidence estimation	Resampling times
CM1 inv - CM1 data/ ASMC-DREAM	[-652.03; -641.02]	-679.48	1
CM1 inv - CM1 data/ MCMC-DREAM	[-666.07; -636.71]	-678.39(LM)	-
CM1 inv - CM1 data/ ASMC-Gauss	[-654.79; -640.65]	-679.80	2
CM2 inv - CM2 data/ ASMC-DREAM	[-628.60; -603.91]	-671.18	5
CM2 inv - CM2 data/ MCMC-DREAM	[-682.90; -612.23]	-	-
CM2 inv - CM2 data/ ASMC-Gauss	[-638.64; -611.15]	-671.49	3
CM1 inv - CM2 data/ ASMC-DREAM	[-1086.42;-1063.34]	-1115.76	5
CM2 inv - CM1 data/ ASMC-DREAM	[-831.70; -795.19]	-919.17	9



Figure 10. Natural log-evidence estimations for ten replications of the ASMC-DREAM algorithm applied to CM1 using K = 1, 3, 5, 10, 20 iterations per intermediate distribution, where each colored point denotes a given replication. The gray dashed line represents the mean of the K = 20 replications and the black stars the corresponding mean for each K.

of strong correlations between the particles. The impact of resampling on the variance estimation 475 in equation 15 is primarily embodied in the sum involving the Eve indices. For smaller K, more 476 resampling is needed and the number of remaining Eve indices are smaller. Figure 11 illustrates 477 the evolution of the Eve indices  $E_t^i$  for K = 1 and K = 5 as the CM1  $\alpha$ -sequence progresses. 478 Of the original 40 Eve indices, there are at the end only 3 and 8 Eve indices surviving for K = 1479 and K = 5, respectively. For K = 20, there are 15 surviving Eve indeces. The larger the number 480 of surviving Eve indices, the less is the risk of mode collapse in which the ASMC algorithm only 481 explore a small part of the posterior distribution. This basically implies that the higher-quality 482 estimates are obtained by using larger K or  $CESS_{op}$ , but this comes at the cost of an increasing 483 number of forward simulations. Table 3 shows the relative standard deviation obtained with Monte 484 Carlo replication and the single ASMC run estimates. For CM1, the relative standard deviations 485 calculated with both estimators are similar for K = 10 and K = 20 suggesting that equations 486 15 and 16 may provide high-quality uncertainty estimates for long-enough ASMC runs. For small 487 K, we observe significant underestimation of the relative standard deviations. For K = 1, the sin-488 gle ASMC estimation is three times smaller than those obtained by Monte Carlo replication. Why 489

does the single-run ASMC uncertainty estimation work well for large K, but not for small ones? To 490 shed some light on this question, we present in Figure 12 the evolution of the difference between 491 the weighted mean of the 40 particles' likelihoods  $\hat{p}(\mathbf{y}|\boldsymbol{\theta})$  and the target log-likelihood calculated 492 with the noise realization  $p_n(\mathbf{y}|\boldsymbol{\theta})$ , both raised to the power of the corresponding  $\alpha$  with the differ-493 ences expressed in logarithmic units, that is,  $\Delta log[p(\mathbf{y}|\boldsymbol{\theta})^{\alpha}] = log[\hat{p}(\mathbf{y}|\boldsymbol{\theta})^{\alpha}] - log[p_n(\mathbf{y}|\boldsymbol{\theta})^{\alpha}]$ . This 494 difference is shown for the ten replications and for the different K-values considered. In addition, 495 Table 3 shows the variance and the root-mean-square error (RMSE) for the last states ( $\alpha = 1$ ) 496  $\Delta log[p(\mathbf{y}|\boldsymbol{\theta}])$  that decrease with increasing K. We observe in Figure 12 that when K decreases, 497 the trajectories becomes more separate and show more auto-correlation. At K = 20 and K = 10498 for which the single-ASMC estimates worked well, we observe that the trajectories overlap and 499 cross each other, thereby, suggesting that the information content of one individual ASMC run is 500 not so much different than another. In contrast, for K = 1 (Figure 12a) the mean trajectories tend 501 to be more separated from each other suggesting that they sample slightly different posteriors. The 502 Monte Carlo replications account for these differences between individual ASMC runs, while this 503 is impossible when considering estimates from a single ASMC run. This suggests then that the 504 single-run evidence estimator should only be trusted when performing a sufficient number of K505 iterations, thereby, ensuring that the approximations of the intermediate distributions for different 506 ASMC runs are small. In practice, this suggests that it is useful to run at least two ASMC runs and 507 to ensure that the weighted mean-likelihoods of their particles are similar and tend to cross mul-508 tiple times during the ASMC runs. If this is not the case, our results suggest that the uncertainty 509 estimation of the evidence obtained from one ASMC run is too small. 510

This finding is also supported by the CM2 estimations in Table 3. This is clearly a more challenging conceptual model, where the K used for the ASMC runs was three times higher than for CM1. Even if the single-run uncertainty estimations decrease consistently when increasing K, the values are too low compared to those of Monte Carlo replication. This suggests that K was not large enough to trust the single-run estimator. This is also reflected in the higher variance and the RMSE of the likelihood difference compared to CM1. This suggests that either Monte Carlo repli-



Figure 11. Eve index evolution vs.  $\alpha$ -sequence evolution for (a) K = 1 and (b) K = 5. The increasing opacity indicates superposition, that is, replication of specific Eve indices for different particles.

- $_{517}$  cations are needed to obtain an accurate error estimator or K should be increased to improve the
- <sup>518</sup> reliability of the single-run estimator.

# 519 4 DISCUSSION

- <sup>520</sup> Our results suggest that ASMC can provide accurate approximations of posterior PDFs for chal-
- <sup>521</sup> lenging inverse problems for which state-of-the-art adaptive MCMC fails to converge when con-

**Table 3.** Relative standard deviation of evidence estimations obtained with ASMC-DREAM using different K iterations per intermediate distribution. Results are shown for estimates based on a single run (equations 16 and 15) and by ten replications for CM1 and five replications for CM2 of the ASMC algorithm. Variance and root-mean-square error (RMSE) of the difference between the average log-likelihoods and the target (noise) log-likelihood are shown for the replications.

K	$\sigma_r$ [single run]	$\sigma_r$ [replications]	$\sigma^2(\Delta log[p(y \boldsymbol{\theta})])$	$RMSE(\Delta log[p(y \boldsymbol{\theta})])$					
	CM1								
1	0.62	1.72	1.70	2.99					
3	0.42	0.66	1.42	1.91					
5	0.35	0.50	0.62	1.84					
10	0.29	0.27	0.67	1.14					
20	0.21	0.20	0.69	1.47					
			CM2						
5	0.47	1.92	8.45	43.34					
10	0.40	1.56	4.16	21.59					
20	0.38	1.02	3.66	13.89					
40	0.36	1.52	5.06	7.70					
60	0.33	1.22	6.36	2.46					



Figure 12. Evolution of the difference between the weighted mean log-likelihood  $\hat{p}(\mathbf{y}|\boldsymbol{\theta})$  and the target log-likelihood calculated with the noise realization  $p_n(\mathbf{y}|\boldsymbol{\theta})$  raised to  $\alpha$ , where each color represents one replication, for (a) K = 1, (b) K = 3, (c) K = 5 (d) K = 10, and (e) K = 20

sidering a similar number of forward simulations (Figure 8). Furthermore, ASMC is very well suited for parallel computation, which is less the case for most MCMC methods. A general recommendation for practical applications is that the algorithmic variables K and  $CESS_{op}$  in Algorithm 1 are chosen sufficiently large to ensure that the weighted-mean likelihood of the particles is close to the target likelihood during the ASMC run (Figure 12). Clearly, if the total number of forward simulations are insufficient, the ASMC algorithm fails in sampling posterior realizations of high likelihood for most particles. This leads to an impoverished particle approximation of the posterior
 PDF as evidenced by few surviving Eve indices (Figure 11) and mode collapse.

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A similar argument holds for the evidence estimation. ASMC provides an unbiased estimation, 530 as shown for the high-noise setting example (section 3.5). However, the evidence estimation pro-531 cedure will only be reliable if the particles approximate the target power posteriors well enough. 532 In addition, too low K and  $CESS_{op}$  lead to frequent resampling that increases the estimation 533 variance. Our results also suggest that error approximations based on single ASMC runs (eqs. 15 534 and 16) are too optimistic in such settings, but reliable for sufficiently long ASMC runs (Table 3). 535 We also note that the relative standard deviations of the evidence estimates (Figure 10) are several 536 orders of magnitude smaller than the evidences obtained for the consistent and inconsistent prior 537 models (Table 2). 538

Providing practical recommendations for parameter settings away from easily-recognizable 539 degenerate conditions is challenging. Of course, the larger the N the better, as the particle ap-540 proximation of the parameter space will be improved. Our choice of N = 40 was dictated by the 541 number of forward runs we could perform in parallel on one compute node, while much larger 542 values are possible on modern computational architectures. An important point is how well the 543 posterior can be described by a weighted average of N particles. The complexity of the posterior 544 distribution depends on several factors like the dimension of the parameter space, the physics, the 545 number and type of data, and the experimental design. Consequently, a much larger number of par-546 ticles might be needed in challenging high-dimensional settings with strong parameter correlations 547 or for problems with multi-modal posterior PDFs. In agreement with Neal (2001), we recommend 548 distributing the total number of forward runs for each ASMC particle by favouring a large number 549 of intermediate distributions over larger K. In practice, we typically first choose a suitably large 550  $CESS_{op}$  and then vary K. In contrast to K, the influence of  $CESS_{op}$  on the total number of for-551 ward simulations is non-linear and difficult to predict before running the algorithm. The trial tests 552 in this study suggest that  $CESS_{op}$  needs to be larger than 0.99N, for our considered ranges of 553 K, in order to reach the target misfit and build a smooth  $\alpha$ -sequence. After fixing  $CESS_{op}$ , one 554 can then first run the ASMC with an initially small K before re-running it with a twice as large 555

value. If the difference between the resulting evidence estimates for these two choices of K are much smaller than the computed evidences for competing conceptual models, and if the inferred posteriors are similar, then this choice of K is probably sufficient. If important differences are observed between the ASMC runs obtained for the different K, then one needs to further double K, and so on. Finally, the proposal scale  $\epsilon$  needs to be initialized with a high enough value such that the initial acceptance rate is above  $AR_{min}$ . After this, the automatic rescaling of this parameter ensures high-quality estimates regardless of the model proposal scheme.

The observed relative insensitivity of the ASMC results to the model proposal type (Figures 4 563 and 5) is noteworthy, as the MCMC results (Figure 6) are highly sensitive to this choice. CM1 and 564 CM2 present different levels of complexity. For CM1, MCMC-DREAM achieves convergence 565 without difficulty (Fig. 6i), while this is far from being the case for MCMC-Gauss (Figure 6j). 566 For CM2, both MCMC approaches fail (Figures 6k and 1), while ASMC-DREAM and ASMC-567 Gauss perform similarly well for both CM1 and CM2 (Figures 4 and 5). The underlying reason 568 for the success of ASMC and its insensitivity to the proposal mechanism is likely found due to 569 the following factors. On the one hand, the adaptive scaling of the proposals (e.g., Figure 4c) and 570 the tempering (e.g., Figure 4d) allow the particles to more easily move away from local minima, 571 while resampling, on the other hand, gives priority to the high-likelihood regions (e.g., Figure 4h). 572 Clearly, no such tuning of the proposal scale is possible when using MCMC as it violates detailed 573 balance conditions. We stress that the comparisons made herein are with MCMC algorithms run-574 ning at a unitary temperature, while parallel tempering-based MCMC methods might not have 575 these problems (Sambridge, 2014). 576

<sup>577</sup> The presented ASMC method share similarities to other approaches for evidence estimation. <sup>578</sup> Nested Sampling (Skilling, 2004) reduces the evidence multidimensional integral to sampling of <sup>579</sup> a one-dimensional integral over prior mass elements, using an increasing constraint on the log-<sup>580</sup> likelihood lower bound. Other methods rely on MCMC sampling using power posteriors. For <sup>581</sup> instance, thermodynamic integration (TIE) (Gelman & Meng, 1998), also called path sampling, <sup>582</sup> reduces the evidence computation to a one-dimensional integral of the expectation of the likeli-<sup>583</sup> hood over  $\alpha$ . Zeng et al. (2018) shows that TIE performs better than nested sampling in terms

of accuracy and stability. Stepping Stone Sampling (SS) (Xie et al., 2011) also rely on power-584 posteriors but improves in accuracy compared with TIE by formulating the evidence estimation by 585 the product of ratios of intermediate normalizing constants, that is, similarly to AIS and ASMC. An 586 important practical difference is that SS is often performed in parallel by running multiple MCMC 587 runs targeting different power posteriors (Brunetti et al., 2019). Since each chain starts from the 588 prior, the total computational cost is high, and perhaps more importantly, there is no solution to 589 deal with MCMC chains for  $\alpha$  close to one that do not converge (as in our MCMC trials with both 590 MCMC-Gauss and MCMC-DREAM for CM2). This latter problem can be circumvented by run-591 ning the SS algorithm sequentially using a similar tempering sequence as for ASMC. However, the 592  $\alpha$ -sequence needs to be pre-defined, while ASMC allows for adaptive tuning. Even if not presented 593 here, we stress that the improvements offered by ASMC over AIS are drastic. Despite extensive 594 testing and tuning of AIS parameters, we were unable to match the performance of ASMC. 595

# 596 5 CONCLUSIONS

This study demonstrates that adaptive sequential Monte Carlo (ASMC) is a powerful method to 597 approximate the posterior PDF and estimate the evidence in non-linear geophysical inverse prob-598 lems. Crosshole GPR examples in which complex geological priors are parameterized through 599 deep generative networks are used for demonstration purposes, but the method is of wide appli-600 cability. ASMC is robust with respect to the type of model proposals used and to algorithmic 601 settings, implying a comparatively low user effort required for tuning the algorithm for a given ap-602 plication. ASMC is particularly useful for moderately to strongly non-linear inverse problems and 603 for multi-modal distributions, where targeting the posterior distribution with MCMC algorithms 604 may result in poor convergence. For the considered examples, ASMC outperforms state-of-the-art 605 adaptive MCMC in estimating posterior PDFs. The major advantage of ASMC compared with 606 MCMC in a Bayesian model selection context is that it provides straightforward computation of 607 the evidence. Reliable uncertainty estimation of evidence estimates is possible from single ASMC 608 runs, provided that they are long enough. We hope that this study will stimulate further adaptations 609 of sequential Monte Carlo in a geophysical context, and more specifically, lead researchers to the 610

adaptation of ASMC when confronted with challenging inference problems and model selection tasks.

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